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Is the Hawking's scenarios of black holes evaporation the only possible solution?

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The author is considering two possible scenarios of black holes evaporation. First one, coincide with well known S Hawking's (1974, 1975) scenarios, according to which actually black holes (created after the Big Bang, perhaps) with $M < (4-7) \cdot 10^{14} g$ should evaporate, while another, which would take into account the occurrence of mass particles bound states (see [1]), especially Bose mass particles, of which the Higgs boson is of special interest. The second scenarios suppose a concurrence between the Hawking process and Bose mass particles exponentially fast accumulation with a rate of Black Hole's mass evolution:

$$M = M_0 \left(1 + \frac{m^2 M_0^2 t}{2\pi(1+s)^2} \right)^{-1/2}, \quad M_0 = \text{initial Black Hole's mass, } s = \text{emitted particles spin Eq. (1)}$$

When the time is going to the mass of a Black Hole is going to 0. The time of diminishing by a half of the initial mass of a Black Hole is:

$$t_{1/2} = \frac{14\pi(1+s)^2}{m^2 M_0^2} \text{ Eq. (2)}$$

which corresponds to $\sim 54700 \text{ sec} = 15.19 \text{ hours}$ for a Black Hole of mass $M_0 \sim 2.25 \cdot 10^{11} g$ (which is nearly 1 mil tones in weight). This would occur due to generation of a Higgs boson ($s=0$) with mass $m_{Higgs} = 170 \text{ GeV} = 0.17(7) \cdot 10^{-20} g$. If the Higgs boson mass is 125 GeV, the time of diminishing by a half of the Black Hole's mass would be $\sim 97 \text{ sec}$. For , when the innermost stable classical (quasi-classical) and quantum mechanical orbit still exist $t_{1/2} \cong 5 \cdot 10^8 s \sim 15.8 y$.

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