# On The Special Curves in <br> Minkowski 4 Spacetime 

Gül Güner<br>Karadeniz Technical University<br>Department of Mathematics<br>Trabzon, Turkey<br>gguner@ktu.edu.tr<br>\section*{F. Nejat Ekmekci}<br>Ankara University<br>Department of Mathematics<br>Ankara, Turkey<br>ekmekci@science.ankara.edu.tr


#### Abstract

In [1], we gave a method for constructing Bertrand curves from the spherical curves in 3 dimensional Minkowski space. In this work, we construct the Bertrand curves corresponding to a spacelike geodesic and a null helix in Minkowski 4 spacetime.


## Mathematics Subject Classification: 53A35

Keywords: Bertrand curve, Minkowski space, Null helix, Spacelike geodesic.

## 1 Preliminary Notes

In this section, we give basic notions of spacelike and null curves in Minkowski 4 -space (see [2], [3] and [6]). Let $\mathbb{R}^{4}=\left\{\left(x_{1}, x_{2}, x_{3}, x_{4}\right): x_{1}, x_{2}, x_{3}, x_{4} \in \mathbb{R}\right\}$ be a 4 -dimensional vector space. For any vectors $x=\left(x_{1}, x_{2}, x_{3}, x_{4}\right), y=$ $\left(y_{1}, y_{2}, y_{3}, y_{4}\right)$ in $\mathbb{R}^{4}$, the pseudo scalar product of $x$ and $y$ is defined to be $\langle x, y\rangle=-x_{1} y_{1}+x_{2} y_{2}+x_{3} y_{3}+x_{4} y_{4}$. We call ( $\left.\mathbb{R}^{4},\langle\rangle,\right)$ a Minkowski 4 -space. We write $\mathbb{R}_{1}^{4}$ instead of $\left(\mathbb{R}^{4},\langle\rangle,\right)$. We say that a non-zero vector $x \in \mathbb{R}_{1}^{4}$ is spacelike, lightlike (null) or timelike if $\langle x, x\rangle>0,\langle x, x\rangle=0$ or $\langle x, x\rangle<0$ respectively. The norm of the vector $x \in \mathbb{R}_{1}^{4}$ is defined by $\|x\|=\sqrt{|\langle x, x\rangle|}$. For a vector $v \in \mathbb{R}_{1}^{4}$ and a real number $c$, we define a hyperplane with pseudo normal $v$ by
$H P(v, c)=\left\{x \in \mathbb{R}_{1}^{4}:\langle x, v\rangle=c\right\}$. We call $H P(v, c)$ a spacelike hyperplane, a timelike hyperplane or a lightlike hyperplane if $v$ is timelike, spacelike or lightlike respectively. We also define de Sitter 3 -space by $S_{1}^{3}=\left\{x \in \mathbb{R}_{1}^{4}\right.$ : $\langle x, x\rangle=1\}$. For any $x=\left(x_{1}, x_{2}, x_{3}, x_{4}\right), y=\left(y_{1}, y_{2}, y_{3}, y_{4}\right), z=\left(z_{1}, z_{2}, z_{3}, z_{4}\right)$ in $\mathbb{R}_{1}^{4}$, we define a vector

$$
x \wedge y \wedge z=\left|\begin{array}{cccc}
-e_{1} & e_{2} & e_{3} & e_{4} \\
x_{1} & x_{2} & x_{3} & x_{4} \\
y_{1} & y_{2} & y_{3} & y_{4} \\
z_{1} & z_{2} & z_{3} & z_{4}
\end{array}\right|
$$

where $\left(e_{1}, e_{2}, e_{3}, e_{4}\right)$ is the canonical basis of $\mathbb{R}_{1}^{4}$. We can easily show that $\langle a,(x \wedge y \wedge z)\rangle=\operatorname{det}(a, x, y, z)$.

Let $\gamma: I \longrightarrow S_{1}^{3}$ be a regular curve. We say that a regular curve $\gamma$ is spacelike, timelike or null respectively, if $\gamma^{\prime}(t)$ is spacelike, timelike or null at any $t \in I$, where $\gamma^{\prime}=d \gamma / d t$. Now we describe the explicit differential geometry on spacelike and null curves in $S_{1}^{3}$.

Let $\gamma$ be a spacelike regular curve, we can reparametrise $\gamma$ by the arclength $s=s(t)$. Hence, we may assume that $\gamma(s)$ is a unit speed curve. So we have the tangent vector $t(s)=\gamma^{\prime}(s)$ with $\|t(s)\|=1$. In the case when $\left\langle t^{\prime}(s), t^{\prime}(s)\right\rangle \neq 1$, we have a unit vector $n(s)=\frac{t^{\prime}(s)-\gamma(s)}{\left\|t^{\prime}(s)-\gamma(s)\right\|}$. Moreover, define $e(s)=\gamma(s) \wedge$ $t(s) \wedge n(s)$, then we have a pseudo orthonormal frame $\{\gamma(s), t(s), n(s), e(s)\}$ of $\mathbb{R}_{1}^{4}$ along $\gamma$. By the standard arguments, we can show the following FrenetSerret type formulae: Under the assumption that $\left\langle t^{\prime}(s), t^{\prime}(s)\right\rangle \neq 1$,

$$
\begin{align*}
\gamma^{\prime}(s) & =t(s) \\
t^{\prime}(s) & =-\gamma(s)+\kappa_{g}(s) n(s) \\
n^{\prime}(s) & =\kappa_{g}(s) \delta(\gamma(s)) t(s)+\tau_{g}(s) e(s) \\
e^{\prime}(s) & =\tau_{g}(s) n(s) \tag{1}
\end{align*}
$$

where $\delta(\gamma(s))=-\operatorname{sign}(n(s))$,

$$
\begin{aligned}
\kappa_{g}(s) & =\left\|t^{\prime}(s)+\gamma(s)\right\| \\
\tau_{g}(s) & =\frac{\delta(\gamma(s))}{\kappa_{g}^{2}(s)} \operatorname{det}\left(\gamma(s), \gamma^{\prime}(s), \gamma^{\prime \prime}(s), \gamma^{\prime \prime \prime}(s)\right)
\end{aligned}
$$

Now let $\gamma: I \longrightarrow S_{1}^{3}$ be a null curve. We will assume, in the sequel, that the null curve we consider has no points at which the acceleration vector is null. Hence $\left\langle\gamma^{\prime \prime}(t), \gamma^{\prime \prime}(t)\right\rangle$ is never zero. We say that a null curve $\gamma(t)$ in $\mathbb{R}_{1}^{4}$ is parametrized by the pseudo-arc if $\left\langle\gamma^{\prime \prime}(t), \gamma^{\prime \prime}(t)\right\rangle=1$. If a null curve satisfies $\left\langle\gamma^{\prime \prime}(t), \gamma^{\prime \prime}(t)\right\rangle \neq 0$, then $\left\langle\gamma^{\prime \prime}(t), \gamma^{\prime \prime}(t)\right\rangle>0$, and

$$
u(t)=\int_{t_{0}}^{t}\left\langle\gamma^{\prime \prime}(t), \gamma^{\prime \prime}(t)\right\rangle^{1 / 4} d t
$$

becomes the pseudo-arc parameter.
A null curve $\gamma(t)$ in $\mathbb{R}_{1}^{4}$ with $\left\langle\gamma^{\prime \prime}(t), \gamma^{\prime \prime}(t)\right\rangle \neq 0$ is a Cartan curve if $\left\{\gamma^{\prime}(t), \gamma^{\prime \prime}(t), \gamma^{\prime \prime \prime}(t)\right\}$ is linearly independent for any $t$. For a Cartan curve $\gamma(t)$ in $\mathbb{R}_{1}^{4}$ with pseudo-arc parameter $t$, there exists a pseudo orthonormal basis $\left\{L, N, W_{1}, W_{2}\right\}$ such that

$$
\begin{align*}
L & =\gamma^{\prime} \\
L^{\prime} & =W_{1} \\
N^{\prime} & =-\gamma+k_{1} W_{1}+k_{2} W_{2} \\
W_{1}^{\prime} & =-k_{1} L-N \\
W_{2}^{\prime} & =-k_{2} L \tag{2}
\end{align*}
$$

where $\langle L, N\rangle=1,\left\langle L, W_{1}\right\rangle=\left\langle L, W_{2}\right\rangle=\left\langle N, W_{1}\right\rangle=\left\langle N, W_{2}\right\rangle=\left\langle W_{1}, W_{2}\right\rangle=$ 0 . We call $\left\{L, N, W_{1}, W_{2}\right\}$ as the Cartan frame and $\left\{k_{1}, k_{2}\right\}$ as the Cartan curvatures of $\gamma$. Since the Cartan frame is unique up to orientation, the number of the Cartan curvatures is minimum and the Cartan curvatures are invariant under Lorentz transformations, the set $\left\{L, N, W_{1}, W_{2}, k_{1}, k_{2}\right\}$ corresponds to the Frenet apparatus of a space curve. A direct computation shows that the values of the Cartan curvatures are

$$
\begin{align*}
& k_{1}=\frac{1}{2 a^{2}}\left(\left\langle\gamma^{\prime \prime \prime}, \gamma^{\prime \prime \prime}\right\rangle+2 a a^{\prime \prime}-4\left(a^{\prime}\right)^{2}\right) \\
& k_{2}=-\frac{1}{a^{4}} \operatorname{det}\left(\gamma^{\prime}, \gamma^{\prime \prime}, \gamma^{\prime \prime \prime}, \gamma^{(4)}\right) \tag{3}
\end{align*}
$$

Theorem 1.1 Let $\gamma(t)$ in $\mathbb{R}_{1}^{4}$ be a Cartan curve. Then $\gamma$ is a pseudospherical curve iff $k_{2}$ is a nonzero constant.

Theorem 1.2 A Cartan curve $\gamma(t)$ in $\mathbb{R}_{1}^{4}$ fully lies on a pseudo-sphere iff there exists a fixed point $A$ such that for each $t \in I,\left\langle A-\gamma(t), \gamma^{\prime}(t)\right\rangle=0$.

## 2 Bertrand Curve Corresponding to A Spacelike Geodesic on $\mathrm{S}_{1}^{3}$

Theorem 2.1 Let $\gamma$ be a spacelike geodesic curve on $S_{1}^{3}$. Then,

$$
\tilde{\gamma}(s)=a \int \gamma(v) d v+a \operatorname{coth} \theta \int e(v) d v+c
$$

is a Bertrand curve where a and $\theta$ are constant numbers, $c$ is a constant vector.
Proof. We will use the frame $\{\gamma(s), t(s), n(s), e(s)\}$ of $\gamma$ given in the previous section. In this frame, let we choose $e(s)$ as a timelike vector (If $e(s)$ is a
spacelike vector, the proof is similar). Hence $n(s)$ is spacelike and $\delta(\gamma(s))=$ -1 . Using the equation (1), we can easily calculate that

$$
\begin{aligned}
\tilde{\gamma}^{\prime}(s)= & a[\gamma(s)+\operatorname{coth} \theta e(s)] \\
\tilde{\gamma}^{\prime \prime}(s)= & a\left[t(s)+\operatorname{coth} \theta \tau_{g}(s) n(s)\right] \\
\tilde{\gamma}^{\prime \prime \prime}(s)= & a\left[-\gamma(s)+\delta(\gamma(s)) \kappa_{g}(s) \tau_{g}(s) t(s)\right. \\
& \left.+\left(\kappa_{g}(s)+\operatorname{coth} \theta \tau_{g}^{\prime}(s)\right) n(s)+\operatorname{coth} \theta \tau_{g}^{2}(s) e(s)\right]
\end{aligned}
$$

Since $\left\langle\tilde{\gamma}^{\prime}(s), \tilde{\gamma}^{\prime}(s)\right\rangle=-\frac{a^{2}}{\sinh ^{2} \theta}$, the curve $\tilde{\gamma}$ is timelike. If we calculate the first and second curvatures of $\tilde{\gamma}$ by using the equations in [8], we have

$$
\begin{aligned}
\kappa(s) & =\frac{\sinh ^{2} \theta \sqrt{1+\operatorname{coth}^{2} \theta \tau_{g}^{2}}}{a} \\
\tau(s) & =\frac{A \sinh \theta}{a \sqrt{1+\operatorname{coth}^{2} \theta \tau_{g}^{2}}}
\end{aligned}
$$

where $A=\sqrt{\cosh ^{2} \theta\left(\tau_{g}^{2}+1\right)^{2}-\kappa_{g}^{2}\left(1+\operatorname{coth}^{2} \theta \tau_{g}^{2}\right)}$. Since $\tau_{g}$ and $\kappa_{g}$ are constants, we can choose $\beta=\frac{-a \sinh \theta \sqrt{1+\operatorname{coth}^{2} \theta \tau_{g}^{2}}}{A}$ and $\alpha=\frac{a \operatorname{coth}^{2} \theta}{\sqrt{1+\operatorname{coth}^{2} \theta \tau_{g}^{2}}}$, then we have $\alpha \kappa+\beta \tau=1$. Hence $\tilde{\gamma}$ is a Bertrand curve.

## 3 Bertrand Curve Corresponding to A Null Helix on $S_{1}^{3}$

Theorem 3.1 Let $\gamma$ be a null helix on $S_{1}^{3}$. Then,

$$
\tilde{\gamma}(s)=a \int L(v) d v+a \operatorname{coth} \theta \int W_{2}(v) d v+c
$$

is a Bertrand curve where a and $\theta$ are constant numbers, $c$ is a constant vector.

## Proof.

$$
\begin{aligned}
\tilde{\gamma}^{\prime}(t) & =a\left[L(s)+\operatorname{coth} \theta W_{2}(t)\right] \\
\tilde{\gamma}^{\prime \prime}(t) & =a\left[1-\operatorname{coth} \theta k_{2}\right] W_{1}(t) \\
\tilde{\gamma}^{\prime \prime \prime}(t) & =a\left[k_{1}(\operatorname{coth} \theta-1) L(t)-\left(1-\operatorname{coth} \theta k_{2}\right) N(t)\right]
\end{aligned}
$$

Since $\left\langle\tilde{\gamma}^{\prime}(t), \tilde{\gamma}^{\prime}(t)\right\rangle=a^{2} \operatorname{coth}^{2} \theta$, the curve $\tilde{\gamma}$ is spacelike. If we calculate the first and second curvatures of $\tilde{\gamma}$, we have

$$
\begin{aligned}
& \kappa(t)=\frac{\left(1-\operatorname{coth} \theta k_{2}\right)}{a \operatorname{coth}^{2} \theta} \\
& \tau(t)=\frac{\sqrt{k_{1}^{2} \cosh ^{2} \theta-1}}{\cosh \theta}
\end{aligned}
$$

Since $k_{1}$ and $k_{2}$ are constants, we can choose $\beta=-\frac{\cosh ^{3} \theta}{\sqrt{k_{1}^{2} \cosh ^{2} \theta-1}}$ and $\alpha=\frac{a \cosh ^{2} \theta}{\left(1-\operatorname{coth} \theta k_{2}\right)}$, then we have $\alpha \kappa+\beta \tau=1$. Hence $\tilde{\gamma}$ is a Bertrand curve.

## References

[1] Güner, G., Ekmekci, N., On the Spherical curves and Bertrand curves in Minkowski-3 space, J. Math. Comput. Sci. 2, 4, 2012, 898-906.
[2] Çöken, A. C., Çiftçi Ü., On The Cartan Curvatures of a Null Curve in Minkowski Spacetime, Geometriae Dedicata, 114, 2005, 71-78.
[3] Ferrandez, A., Gimenez, A., Lucas, P., Characterization of null curves in Lorentz-Minkowski spaces, Publicaciones de la RSME, 3, 2001, 221-226.
[4] Liu, H., Curves in the Lightlike Cone, Contributions to Algebra and Geometry, 1, 2004, 291-303.
[5] İlarslan, K., Nesovic, E., Some Characterizations of Null, Pseudo Null and Partially Null Rectifying Curves in Minkowski Space-Time, Taiwanese Journal of Mathematics, 5, 2008, 1035-1044.
[6] Fusho, T., Izumiya, S., Lightlike surfaces of spacelike curves in de Sitter 3 -space, 2006.
[7] Matsuda, H., Yorozu, S., Notes on Bertrand curves, Yokohama Mathematical Journal, 50, 2003, 41-58.
[8] Yılmaz, S., Turgut, M., On the Differential Geometry of the Curves in Minkowski Space-time II, Int. J. Contemp. Math. Sciences, 3, 2, 2009.

Received: October, 2014

